Course: M.Sc (Physics)

| Semester | 4 |
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| Paper Number | 17 (MPHC4403) |
| Paper Title | Galaxies and Particle Astrophysics |
| No. of Credits | 6 |
| Course description/objective | The aim is also to enable a student to visualize the geometric structure of space-time represented by a black hole solution of Einstein's equation and the trajectories of photons in such space-times. It also aims to teach the students that accretion is one of the most important energy generation mechanisms. This course aims to educate the students about some major components of the universe like the galaxies and interstellar medium, about their classification and the physical principles involved in their formation and evolution. Group B: In this part of the course, the students are taught canonical quantization of the real scalar field, the complex scalar field, the electromagnetic field and the Dirac field. They are also introduced to the concept of spontaneous symmetry breaking and the generation of mass by the Higgs' mechanism |
| | The Objective of this course is to learn about the history and evolution of the idea of using cosmic rays as astrophysical probes. To this ends the course takes the students through possible scenarios of cosmic ray generation, acceleration, interaction with terrestrial atmosphere and their detection methods. It also provides an introduction to candidate dark matter objects and experiment along with a brief description of gravitational wave phenomena. |
| | <u>Group A</u> |
| | <u>Compact Objects:</u> |
| Course Outcome | CO1: To understand the basic mechanisms involved in the formation of compact objects like neutron stars, black holes, x-ray sources, pulsars, etc. CO2: To gather an in-depth knowledge about the mechanism of gravitational collapse leading to the formation of black holes, and the significance of trapped surfaces and horizons. |
| | CO3: To apply this knowledge to understand the coordinate transformations used to handle a singularity and the corresponding change in the nature of light cones, so as to understand the global properties and causal structure of space-times. CO4: To learn about the structure and evolution of white dwarfs and neutron stars so as to analyze the conditions of stability of white dwarfs and study the parameters associated with neutron star binaries. |
| | Interstellar Medium and Galaxies: |
| | CO1: Understand the birth and collapse of stars in the interstellar medium CO2: Appreciate that the physical laws that can be applied to the ISM can be extended to intergalactic medium and galaxies CO3: Learn and appreciate galaxies and their evolution as a major component of the structures of the universe. |
| | <u>Group B</u> |
| | Standard Model: |

| CO1: After completing this course, the students should be able to quantize the scalar fields, the electromagnetic field and the Dirac field, canonically. CO2: They should also have obtained a basic understanding of spontaneous symmetry breaking. |
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| Astroparticle Physics: |
| CO1: Develop skill in relativistic dynamics using examples from the physics of |
| interaction of cosmic rays with particles in IGM / Terrestrial atmosphere. |
| CO2: Learn about the operating principles of Cosmic ray detectors and how they extend |
| and contribute to multi-messenger astronomy |
| CO3: Understand the implications of massive neutrino states |
| CO4: Learn about particle candidates of Dark matter |

| | Group A: | |
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| | Compact Objects, Interstellar Medium & Galaxies | [36 Lectures] |
| | Compact Objects lecturesl | [24 |
| | Black holes and compact objects. White Dwarfs and Neutron Stars, Co Sources, Radio Pulsars, Supermassive and other Black Holes, Gamma-Ra | ompact X-Ray y Bursters. [4 lectures] |
| | Collapse of stars and formation of black holes; singularities and trapped horizon; Schwarzschild Black Hole, Tortoise Coordinates and Null Cone Extension, Penrose Diagram – the Conformal Structure of Infinity. Ker Boyer–Lindquist Coordinates, ergosphere; Weyl–Papapetrou form of the uniqueness of the Kerr solution. Dark Energy stars. Black hole accretion disk properties. | surfaces, event s, The Kruskal rr Black Hole; e Kerr Metric, 1 and accretion |
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| Syllabus | White Dwarfs and Classification. Equation of State below the Neutron Structure of White Dwarfs and the Chandrasekhar Mass, Polytropic A Relativistic Instability of White Dwarf Stars, Necessary Condition for S Energy in the Post-Newtonian Limit, GR White Dwarf Instability. | Drıp Densıty, Approximation, Stability, Total |
| | [5 | 5 lectures] |
| | Structure of a Neutron Star, Equations of State beyond Neutron Drip, Drip to Saturation, Nuclear EoS for Dense Neutron Matter. Structur Neutron Stars. Periastron Shift, Shapiro Time Delay in a Binary System. | From Neutron re of Massive |
| | | [5 lectures] |
| | Interstellar Medium and Galaxies | [12 lectures] |
| | Interstellar medium and star formation: Interstellar dust and gas, forma stars, pre-main sequence evolution, initial mass function | ation of proto- |
| | | [5 lectures] |
| | Galaxies: The Milky Way, Distribution of stars, Morphology, Kinen hydrogen(HI 21cm line)Dark matter, Galactic Center, Norn morphological classification, physical characteristics and kinematics Medium. | natics, Neutral mal galaxies: , Intergalactic |

| | | [7 lectures] |
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| | Group B: | |
| | Standard Model and Astroparticle Physics | [36 Lectures] |
| | Standard Model | [24 Lectures] |
| | Standard Model of Particle Interactions, Classical Field theory: Noether's applications. [4 lectures] | Theorem and |
| | Quantum Field Theory: Quantization of the Klein-Gordon, Dirac and Elec Field. [10 lectures] | ctromagnetic |
| | Gauge theories of electromagnetic and weak interactions. | |
| | | [6 lectures] |
| | Spontaneous Symmetry breaking and elementary idea about the Higgs' m | echanism. |
| | | [4 lectures] |
| | Astroparticle Physics [12 | Lectures] |
| | Astroparticle physics in perspective: Relationship to High Energy Physic and Cosmology. Elementary particle processes in galactic, intergalactic a production and transmissions, Physics of particle and radiation detection. | s, Astrophysics nd atmospheric [4 lectures] |
| | Acceleration mechanisms in sunspots, supernovae shocks and accreting b | inaries. [4 lectures] |
| | Primary and secondary cosmic rays: Charged and uncharged compor- gamma, x-ray and gravitational wave astronomy. Secondary cosmic rays: Atmospheric propagation, cosmic rays at the sea level and underground showers (elementary theory). | ents: neutrino, l, Extensive air [4 lectures] |
| | Group A: | |
| References | Compact Objects 1.M. Camenzind, Compact Objects in Astrophysics 2. Stuart L. Shapiro and Saul A. Teukolsky, Black Holes, White Dwarf Stars – The Physics of Compact Objects 3. Sean Carroll, Spacetime and Geometry 4. J. B. Hartle, Gravity – An Introduction to Einstein's General Relativity | s, and Neutron |
| | Interstellar Medium and Galaxies 1. An introduction to the theory of stellar structure and evolution by Dina | Prialnik |

| | (Cambridge University Press) |
|------------|---|
| | 2. Physics of the Interstellar and Intergalactic Medium by B. T. Draine (Princeton Series |
| | in Astrophysics) |
| | 3. The Physics of the Interstellar Medium, J. E. Dyson and D. A. Williams (CRC Press) |
| | 3. Galaxies in the Universe by Sparke and Gallagher, Cambridge University Press |
| | (2007) |
| | 4. An Introduction to Modern Astrophysics by B. W. Carrol and D. A. Ostlie (Pearson, |
| | 2006) |
| | 5. Galactic Astronomy by J. Binney & M. Merrifield (Princeton Series in Astrophysics) |
| | 6. Galactic Dynamics by J. Binney & S. Tremaine (Princeton Series in Astrophysics) |
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| | Group B: |
| | Standard Model |
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| | 1.A first course in quantum field theory by A. Lahiri & P. B. Pal |
| | 2. Quarks and Leptons: an introductory course in modern particle physics by F. Halzen |
| | & A. D. Martin. |
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| | Astroparticle Physics |
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| | 1. Vanessa Cirkel-Bartelt - History of Astroparticle Physics |
| | 2. Claus Grupen : Astroparticle Physics |
| | 3. Donald Perkins: Particle Astrophysics |
| | 4. Claus Grupen : Particle Detectors |
| | [Total: 100 |
| Evaluation | CIA: 10 (Group A) + 10 (Group B) 12 + 10 = 10 |
| | End Semester Examination: 40 (Group A) + 40 (Group B) |